Heat, Light, and Sound Research, Inc.

## BELLHOP3D User Guide

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Heat, Light, and Sound Research, Inc.

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# BELLHOP3D USER GUIDE 

## I. Introduction

BELLHOP3D is a beam tracing model for predicting acoustic pressure fields in ocean environments. It is an extension to 3D environments of the popular BELLHOP model and includes (optionally) horizontal refraction in the lat-long plane. Of course, 3D pressure fields can be calculated by a 2D model simply by running it on a series of radials (bearing lines) from the source. This is the so-called Nx2D or 2.5D approach. However, that approach neglects the refraction of sound energy out of the vertical plane associated with each bearing line. Such out-of-plane effects can be important when there are significant horizontal gradients in the environment. These occur with strong oceanographic features such as nonlinear internal waves, or in areas with strong bathymetric features. This is currently an active area of research.

A very preliminary research version of BELLHOP3D was written in 1985 in FORTRAN. However, it did not allow environmental information (sound speed and bathymetry) to be read in from input files. Instead, the research version required a user-defined analytic function for the sound speed, $c(x, y, z)$, as a function of latitude, longitude, and depth. The research version also did not allow variable bathymetry.

Separately a Matlab conversion of the original research code was done around 2004. The Matlab environment is much easier to develop in; however, we have chosen not to build off of that here, since the Matlab code runs much more slowly (about 50 times slower).

The beam tracing structure leads to a particularly simple algorithm. Several types of beams are implemented including Gaussian and hat-shaped beams, with both geometric and physics-based spreading laws. BELLHOP3D can produce a variety of useful outputs including transmission loss, eigenrays, arrivals, and received time-series. It allows for lat-long variation in the top and bottom boundaries (altimetry and bathymetry), as well as full 3D variation in the sound speed profile. Additional input files allow the specification of directional sources as well as geoacoustic properties for the bounding media. Top and bottom reflection coefficients may also be provided. BELLHOP3D is implemented in Fortran with Matlab wrappers to display the input and output. The code

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conforms closely to standards and can be used on a variety of platforms (Mac, Windows, and Linux). This report describes the code and illustrates its use.

## II. Under the Hood: the BELLHOP algorithm

## A. Introduction: Beam types and overview of beam tracing

BELLHOP3D includes 4 different types of beams:

- Cerveny Beams: We use this term to refer to the original beams derived in the 1982 paper by Cerveny, Popov, and Psencik. These beams evolve roughly as a true propagating beam would; however, the beam field is only accurate near the axis of the ray which can cause problems.
- Geometric Hat-Beams: Much more reliable than the Cerveny beams; recovers conventional ray theory but with the advantages of a beam-tracing algorithm. Good framework for algorithmic testing because any flaws are obvious.
- Geometric Gaussian-Beams: Much improved accuracy relative to the hat beams with some increase in run time.
- Geometric Hat-Beams in Cartesian Coordinates: Improved efficiency compared to the hat beams in ray-centered coordinates. Results are essentially identical to the ray-centered beams.

These last two options should generally be the standard for use with BELLHOP3D; however, we provide the other choices for research purposes.

The initial version of BELLHOP3D has been developed using what we call 'Cerveny' beams. This style of Gaussian beam follows the classic 1983 paper by Cerveny, Popov, and Psencik in which the beams evolve according to physical laws. In other words, they focus and defocus roughly as a real beam in space would do. The reason for starting with this formulation is essentially historical.

In practice, with our 2D version of BELLHOP we had implemented quite a few different types of beams. Originally BELLHOP (2D) also started off with the Cerveny beams. The first change we did was to implement a formulation of the Cerveny beams is Cartesian coordinates rather than raycentered coordinates. Ray-centered coordinates are the arclength, s, and normal distance from the ray, n . Cartesian coordinates are the range coordinate, $r$, and the vertical distance, delta $z$, from the ray. The representation in Cartesian coordinates leads to a much more efficient code for evaluating the contribution of a given beam to the total acoustic field.

The next stage in the evolution of BELLHOP was to implement 'geometric' hat-shaped beams. These beams are not physical per se, but instead spread according to the expansion of a ray-tube. By using a hat-shape (triangle) for each beam one obtains a construction inspired by the finiteelement literature. The result recovers precisely the ray-theoretic result. However, the beam
implementation provides an elegant framework for the implementation and yielded a ray-tracing result that was free of many of the traditional problems of ray models.

Finally, we replaced the hat-shaped beams with Gaussian beams but still implement with a geometric spreading function. Further, we implemented a stop on the minimum width of the beams to eliminate caustics in the ray field.

In summary, the 2D version of BELLHOP has beams implemented in both ray-centered and Cartesian coordinates. It also has beam types that spread physically (Cerveny) or geometrically. The geometric beams are either hat-shaped or Gaussian. Finally, they may optionally have a limit on the minimum width.

The obvious question then is which is the best? The answer is that we generally obtain the best accuracy with geometrically-spreading Gaussian beams that have the minimum-width constraint. However, we also often use the geometrically-spreading hat-shaped as they are faster. They are faster because with the Gaussian beams one gets many contributing beams of a given class at a specific receiver location; with hat-shaped beams, exactly two beams of a given class contribute.

To get the hat-shaped geometric beams to work, one must be extremely careful in the implementation. If the beams are too narrow, then gaps between the beams will show up in the field plots. If they are too wide, then that is also easy to observe in the beam plots. Once they are correctly implemented, then the Gaussian beam version is trivial to implement and smooths out many of the flaws of ray theory.

A further subtlety that is unique to the 3D version of Gaussian beams is that they can be implemented two ways. The traditional form from the literature solves the beam spreading equations twice. The first solution gives the spreading in the vertical plane. The second solution gives the spreading in the horizontal plane. These are called fundamental solutions.

In our original formulation we combined these using a technique that is not well-known, called variously the delta-matrix formulation or compound-matrix formulation. This technique recognizes that all the final quantities of interest to construct the Gaussian beams involve determinants of the fundamental solutions. Rather than solve differential equations for the fundamental solutions, we solve a new set of differential equations for the determinants of the fundamental solutions.

However, to simplify our development and debugging, we have gone back during to a formulation using the fundamental solutions. This allows us to check the spreading behavior of the beams in the various planes.

To illustrate the features of the different beam types we consider a canonical deep water sound speed profile. First, we show in Fig. 1 how a Cerveny beam in BELLHOP3D looks in side view. Note that the beam refracts within the water column and focuses and defocuses due to those
same refractive effects. At the frequency used $(50 \mathrm{~Hz})$ the wavelength is about 30 m so it is difficult to get beams narrower than a few hundred meters. (This is discussed in more detail in the latest edition of Computation Ocean Acoustics.) The fact that the beams are this large is in fact one of the limitations of the method, since the beam field is only accurate near the axis of the beam.


Figure 1: A Cerveny-style beam as first implemented in BELLHOP3D.
Figure 2 shows the new implementation of geometric, hat-shaped beams. In this case, the width of the beams is determined entirely by the spacing of the fan of rays. As we increase the number of rays that are traced, the beams get narrower and narrower. To check the implementation, we compare the result to the geometric, hat-shaped beams in BELLHOP3D, but using the Nx2D option (Figure 3). This is essentially the same as what is in the standard 2D version of BELLHOP and has been tested extensively. We see that there is an excellent correspondence between the two.

Finally, we compute the entire acoustic field by summing up the contributions of a fan of such beams. The comparison of BELLHOP3D with geometric beams using the full 3D option and the $\mathrm{N} \times 2 \mathrm{D}$ option is shown in Figure 4 and is excellent. These results have also been checked again normal mode and wavenumber integration solutions and are of high quality.

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Figure 2: A geometric beam in the full 3D option.


Figure 3: A geometric beam in the Nx2D option.

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Figure 4: Transmission loss calculated with geometric beams with the full 3D option.


Figure 5: Transmission loss calculated using geometric beams in the Nx2D option.

## B. Theory

The extension of Gaussian beams to three dimensions is described by Babich and Popov and requires fairly minor modifications to the 2D algorithm. Let us go through the three steps required for constructing the beam solution. As usual we begin by tracing a set of rays however in the 3D case the rays form a fan over both azimuthal and elevation-declination angles. The ray equations in 3D are given by

$$
\begin{array}{ll}
\frac{d x}{d s}=c \xi(s), & \frac{d \xi}{d s}=-\frac{1}{c^{2}} \frac{d c}{d x} \\
\frac{d y}{d s}=c \eta(s), & \frac{d \eta}{d s}=-\frac{1}{c^{2}} \frac{d c}{d y} \\
\frac{d z}{d s}=c \zeta(s), & \frac{d \zeta}{d s}=-\frac{1}{c^{2}} \frac{d c}{d z}
\end{array}
$$

where $c(x, y, z$,$) is the ocean sound speed (x(s), y(s), z(s))$ is the ray trajectory. This first-order system of ordinary differential equations (ODEs) is integrated using a simple second-order RungeKutta method. The initial conditions prescribe that the rays emanate from the source position ( $\mathrm{x}_{\mathrm{s}}$, $y_{s}, z_{s}$ ) and with take-off angles alpha and beta corresponding to the declination angle and the azimuthal angle of the ray:

$$
\begin{array}{ll}
x(0)=x_{s}, & \xi=\frac{1}{c(0)} \cos \alpha \cos \beta \\
y(0)=y_{s}, & \eta=\frac{1}{c(0)} \cos \alpha \sin \beta \\
z(0)=z_{s}, & \zeta=\frac{1}{c(0)} \sin \alpha
\end{array}
$$

The Gaussian beam is constructed around a central ray and defined in terms of ray-centered coordinates ( $s, m, n$ ) as indicated in the figure. Here s is the arclength along the ray, and $(m, n)$ are normal distances from a field point to the central ray. It is usually not an important issue; however, this coordinate system has a zone of regularity around the central ray where a receiver point has a well-defined ray-centered coordinate. However, for some other receiver points there is more than one normal from the ray to the receiver. For instance if the ray is a circular arc, then radials from the center of the circle are all valid normals.

To be specific, the values of $(m, n)$ are defined as normal distances in the direction of the following two normal vectors to the ray:

$$
\begin{aligned}
\tilde{e}_{m}= & \left(L^{-1}[c \xi \zeta \cos \phi+\eta \sin \phi]\right. \\
& \left.L^{-1}[c \eta \zeta \cos \phi-\xi \sin \phi], c L \cos \phi\right) \\
\tilde{e}_{n}= & \left(L^{-1}[c \xi \zeta \sin \phi-\eta \cos \phi]\right. \\
& \left.L^{-1}[c \eta \zeta \sin \phi+\xi \cos \phi],-c L \cos \phi\right)
\end{aligned}
$$

and $L=\sqrt{\xi^{2}+\eta^{2}}$. (These formulas are derived in Cerveny and Hron. This ray-centered coordinate system $\left(\tilde{t}, \tilde{e}_{m}, \tilde{e}_{n}\right)$ is a rotating trihedral with rotation angle $\phi$ satisfying the differential
equation:

$$
\frac{d \phi}{d s}=\frac{1}{c(s)} \frac{\zeta\left(\eta c_{x}-\xi c_{y}\right)}{\xi^{2}+\eta^{2}}
$$

which must also be integrated along the central ray of the beam. This is one of many ways of calculating the rotating trihedral.

The coordinate system defined by the rotating trihedral is magical. As discussed by Popov (1977) we can use Hamilton's Principle to derive equations that characterize how a ray displaces as we make infinitesimal changes to its original position or take-off angles. These equations are amazingly simple when presented in the coordinate system of the rotating trihedral. Furthermore, the way the ray displaces characterizes the spreading of the ray tube and therefore the intensity along the central ray.

The construction of beams in the neighborhood of the central rays requires the integration of a system of auxilliary equations

$$
\begin{aligned}
\frac{d P}{d s} & =-\frac{1}{c^{2}} V Q \\
\frac{d Q}{d s} & =c P
\end{aligned}
$$

where V is a matrix of curvatures of the sound speed taken in two normal directions:

$$
\begin{gathered}
V=\left[\begin{array}{ll}
c_{n n} & c_{n m} \\
c_{m n} & c_{m m}
\end{array}\right] \\
Q=\left[\begin{array}{ll}
\tilde{q}_{1} & \hat{q}_{1} \\
\tilde{q}_{2} & \hat{q}_{2}
\end{array}\right] \\
P=\left[\begin{array}{ll}
\tilde{p}_{1} & \hat{p}_{1} \\
\tilde{p}_{2} & \hat{p}_{2}
\end{array}\right]
\end{gathered}
$$

The derivatives $c_{n n}, c_{m m}$, etc indicate derivatives in direction of the normal vectors $\tilde{e}_{m}, \tilde{e}_{n}$. We can rewrite these in terms of the $\mathrm{x}, \mathrm{y}, \mathrm{z}$ derivatives of the sound speed as:

$$
\begin{aligned}
c_{n n}= & c_{x x} e_{1 x}^{2}+c_{y y} e_{1 y}^{2}+c_{z z} e_{1 z}^{2}+2 c_{x y} e_{1 x} e_{1 y}+2 c_{x z} e_{1 x} e_{1 z}+2 c_{y z} e_{1 y} e_{1 z} \\
c_{m n}= & c_{x x} e_{1 x} e_{2 x}+c_{y y} e_{1 y} e_{2 y}+c_{z z} e_{1 z} e_{2 z}+ \\
& c_{x y}\left(e_{1 x} e_{2 y}+e_{2 x} e_{1 y}\right)+c_{x z}\left(e_{1 x} e_{2 z}+e_{2 x} e_{1 z}\right)+c_{y z}\left(e_{1 y} e_{2 z}+e_{2 y} e_{1 z}\right) \\
c_{m m}= & c_{x x} e_{2 x}^{2}+c_{y y} e_{2 y}+c_{z z} e_{2 z}^{2}+2 c_{x y} e_{2 x} e_{2 y}+2 c_{x z} e_{2 x} e_{2 z}+2 c_{y z} e_{2 y} e_{2 z}
\end{aligned}
$$

The P-Q differential equation tells us how the ray is perturbed due to a change in the ray initial condition (either by displacing the source position or changing the ray angle). To obtain Gaussian beams, we use the initial conditions:

$$
\left[\begin{array}{ll}
\tilde{p}_{1} & \hat{p}_{1} \\
\tilde{p}_{2} & \hat{p}_{2} \\
\tilde{q}_{1} & \hat{q}_{1} \\
\tilde{q}_{2} & \hat{q}_{2}
\end{array}\right]=\left[\begin{array}{cc}
1 & 0 \\
0 & 1 \\
\epsilon_{1} & 0 \\
0 & \epsilon_{2}
\end{array}\right]
$$

where $\epsilon_{1,2}$ are the beam constants that control the initial beam widths in the two normal directions to the ray. The epsilons are in general complex numbers and their real and imaginary parts allow independent control of both the beam width and the beam curvature (that is, the curvature of its wavefronts).

Once we have integrated these equations along the ray we form a Gaussian beam as:

$$
\begin{gathered}
u_{\text {beam }}(s, m, n)=\frac{1}{\sqrt{|Q(s)|}} e^{i \omega\left[\tau(s)+\frac{1}{2} \vec{q}^{t} \Gamma(s) \vec{q}\right]} \\
\Gamma=P Q^{-1}=\left[\begin{array}{cc}
\tilde{p}_{1} \hat{q}_{2}-\hat{p}_{1} \tilde{q}_{2} & -\tilde{p}_{1} \hat{q}_{1}+\hat{p}_{1} \tilde{q}_{1} \\
\tilde{p}_{2} \hat{q}_{2}-\hat{p}_{2} \tilde{q}_{2} & -\tilde{p}_{2} \hat{q}_{1}+\hat{p}_{2} \tilde{q}_{1}
\end{array}\right] /|Q|
\end{gathered}
$$

The expression for $\boldsymbol{\Gamma}$ has been obtained here by using the cofactor representation of the inverse of Q. In addition $|\mathrm{Q}|$ denotes the determinant of Q . In general the determinant of $Q$ is a complex number that wraps around the origin. To get the correct branch of the square root we need to track these rotations. This produces the so-called KMAH index that accounts for phase changes at caustics.

## C. Geometric Beams

As discussed earlier the geometric beam is designed to have a beam width that expands and contracts in proportion to the ray tube. The hat-shaped beam decays linearly away from the central ray. One may envision a beam that has an amplitude decay in the form of a pyramid; however, since the ray tube does not necessarily have a square cross-section the beam shape is really tetrahedral. The formula for the hat-shaped beam is simply:

$$
u_{\text {beam }}(s, m, n)=\frac{1}{\sqrt{|Q(s)|}} e^{i \omega \tau(s)} \frac{\left[L_{1}(s)-n\right]\left[L_{2}(s)-m\right]}{L_{1}(s) L_{2}(s)}
$$

for $|\mathrm{n}|<\mathrm{L} 1$ and $|\mathrm{m}|<\mathrm{L} 2$. (The beam vanishes outside that tube.) Here L 1 and L 2 are the widths of the ray tube. To get those cross-sectional widths we simply set the initial beam constants, $\epsilon_{1,2}$ to
zero. In that case, the $\mathrm{P}-\mathrm{Q}$ initial conditions imply that we are perturbing the ray in the declination direction (for $\tilde{p}, \tilde{q}_{\text {) or the azimuthal direction (for }} \hat{p}, \hat{q}$ ).

$$
\left[\begin{array}{l}
L_{1}(s) \\
L_{2}(s)
\end{array}\right]=\left[\begin{array}{c}
\delta \alpha \frac{\left|\tilde{q}_{1}(s)\right|}{c(s)} \\
|\cos \alpha| \delta \beta \frac{\left|\hat{q}_{2}(s)\right|}{c(s)}
\end{array}\right]
$$

where $\delta \alpha$ and $\delta \beta$ is the angular spacing between the rays at the source. The hat-shaped beams reproduce the classical ray-theoretic result since the beam intensity is exactly the ray-theoretic result on the central ray and the beam sum produces a linear interpolation for the field in between the rays.

As discussed above, we typically obtain better accuracy by using Gaussian beams but still within the geometric framework. The field at any point is then the sum of contributions from a number of nearby beams, rather than just the two bracketing beams that the hat-shaped beams provide. This integration over many beams smooths out caustics and also allows leakage energy into shadow zones. The formula for the Geometric Gaussian beams is simply:

$$
u_{\text {beam }}(s, m, n)=\frac{1}{\sqrt{|Q(s)|}} e^{i \omega \tau(s)} e^{\frac{-.5\left[\frac{n}{L_{1}(s)}\right]^{2}\left[\frac{m}{L_{2}(s)}\right]^{2}}{L_{1}(s) L_{2}(s)}}
$$

At caustics the beam widths and the determinant of $Q(s)$ vanish. To avoid this singularity we limit the minimum beam width to one wavelength as suggested by Keenan and Burridge.

## D. The Compound Matrix Method

It commonly occurs that one is solving a system of ordinary differential equations with different initial conditions to assemble a final solution that involves determinants of the resulting independent solutions. In this case one can do some simple manipulations to derive a new set of ordinary differential equations that is solved with just one set of initial conditions to yield the determinants directly. This is sometimes called the compound matrix method or the delta-matrix formulation. As we shall see there is typically a redundant equation that can be eliminated leading the to the socalled reduced delta-matrix formulation. The advantage of this approach is that it yields a simpler and faster algorithm.

To derive the new system, we first arrange the original ODE from a matrix to a vector form as follows:

$$
\left[\begin{array}{l}
p_{1} \\
p_{2} \\
q_{1} \\
q_{2}
\end{array}\right]^{\prime}=\left[\begin{array}{cccc}
0 & 0 & c_{n n} & c_{n m} \\
0 & 0 & c_{m n} & c_{m m} \\
c & 0 & 0 & 0 \\
0 & c & 0 & 0
\end{array}\right]\left[\begin{array}{l}
p_{1} \\
p_{2} \\
q_{1} \\
q_{2}
\end{array}\right]
$$

where the prime denotes the derivative with respect to arc length. Next we consider two independent solutions for the p-q vector using tildes for one solution and hats for the other solution. We then assemble those two column vectors into a matrix and construct a new vector consisting of all the principal minors of that matrix. We note that those principal minors correspond to all the elements in the beam field above.

$$
\left[\begin{array}{ll}
\tilde{p}_{1} & \hat{p}_{1} \\
\tilde{p}_{2} & \hat{p}_{2} \\
\tilde{q}_{1} & \hat{q}_{1} \\
\tilde{q}_{2} & \hat{q}_{2}
\end{array}\right] \rightarrow\left[\begin{array}{c}
\tilde{p}_{1} \hat{p}_{2}-\tilde{p}_{2} \hat{p}_{1} \\
\tilde{p}_{1} \hat{q}_{1}-\tilde{q}_{1} \hat{p}_{1} \\
\tilde{p}_{1} \hat{q}_{2}-\tilde{q}_{2} \hat{p}_{1} \\
\tilde{p}_{2} \hat{q}_{1}-\tilde{q}_{1} \hat{p}_{2} \\
\tilde{p}_{2} \hat{q}_{2}-\tilde{q}_{2} \hat{p}_{2} \\
\tilde{q}_{1} \hat{q}_{2}-\tilde{q}_{2} \hat{q}_{1}
\end{array}\right]=\left[\begin{array}{c}
|P| \\
-\Gamma_{12}|Q| \\
+\Gamma_{11}|Q| \\
-\Gamma_{22}|Q| \\
+\Gamma_{21}|Q| \\
|Q|
\end{array}\right]=\left[\begin{array}{c}
\Delta_{p} \\
-h \\
+f \\
-g \\
+h \\
\Delta_{q}
\end{array}\right]=\left[\begin{array}{c}
y_{1} \\
y_{2} \\
y_{3} \\
y_{4} \\
y_{5} \\
y_{6}
\end{array}\right]
$$

Thus, we can write the beam field as:

$$
u_{b e a m}=\frac{1}{\sqrt{|Q|}} e^{i \omega\left[\tau+\frac{1}{2}\left(f n^{2}+2 h m n+g m^{2}\right)\right]}
$$

To get all the terms needed in this new formulation, we require the differential equation for the vector of minors, y . This is obtained by simply differentiating the expression for each principal minor using the product rule. Then from the original ODE we can replace derivatives. The final result is:

$$
\left[\begin{array}{l}
y_{1} \\
y_{2} \\
y_{3} \\
y_{4} \\
y_{5} \\
y_{6}
\end{array}\right]^{\prime}=\left[\begin{array}{cccccc}
0 & -c_{m n} / c^{2} & -c_{m m} / c^{2} & c_{n n} / c^{2} & c_{n m} / c^{2} & 0 \\
0 & 0 & 0 & 0 & 0 & c_{n m} / c^{2} \\
c & 0 & 0 & 0 & 0 & -c_{n n} / c^{2} \\
-c & 0 & 0 & 0 & 0 & c_{m m} / c^{2} \\
0 & 0 & 0 & 0 & 0 & -c_{n m} / c^{2} \\
0 & 0 & c & -c & 0 & 0
\end{array}\right]\left[\begin{array}{l}
y_{1} \\
y_{2} \\
y_{3} \\
y_{4} \\
y_{5} \\
y_{6}
\end{array}\right]
$$

The equation for $\mathrm{y}_{5}$ is just the negative of that for $\mathrm{y}_{2}$, so it can be eliminated. We are left with a firstorder system for five components:

$$
\begin{aligned}
\frac{d Q}{d s} & =\frac{1}{c^{2}}\left(-c_{m m} f+2 c_{m n} h-c_{n n} g\right) \\
\frac{d P}{d s} & =-c(f+g) \\
\frac{d f}{d s} & =c P-\frac{1}{c^{2}} c_{n n} Q \\
\frac{d g}{d s} & =c P-\frac{1}{c^{2}} c_{m m} Q \\
\frac{d h}{d s} & =-\frac{1}{c^{2}} c_{m n} Q
\end{aligned}
$$

Here $\mathrm{C}_{\mathrm{mm}}$ and $\mathrm{C}_{n}$ are second derivatives of the sound speed in the two normal directions $\tilde{e}_{m}, \tilde{e}_{n}$. Once this system is integrated one obtains a representation of the beam as,

$$
u^{b e a m}(s, m, n)=A \sqrt{Q(s)} \exp \left\{-i \omega\left[\tau(s)+\frac{f(s) n^{2}+2 h(s) m n+g(s) m^{2}}{2 Q(s)}\right]\right\}
$$

This representation follows from application of the reduced delta matrix method to the equations given by Babich and Popov. Initial conditions are given forming the same principal minors of the initial conditions from the original $\mathrm{P}-\mathrm{Q}$ equations:

$$
(P, Q, f, g, h)=\left(1, \epsilon_{1} \epsilon_{2}, \epsilon_{2}, \epsilon_{1}, 0\right)
$$

where $\epsilon_{1,2}$ are the (complex) beam constants which characterize the beam width and curvature in directions $\tilde{e}_{m}, \tilde{e}_{n \text { respectively. }}$

Finally, the complex pressure is computed by summing up all of the individual beams with appropriate weighting,

$$
u(r, z)=\sum_{k=1}^{N_{\beta}} \sum_{l=1}^{N_{\alpha}} \delta \beta \delta \alpha \frac{\sqrt{\epsilon_{1} \epsilon_{2}}}{2 \pi{\sqrt{c_{0}}}^{3}} u_{k l}^{b e a m}
$$

E. Beam changes across interfaces

An interface is defined here as a curve across which the sound speed or its derivative is discontinuous. The most important of these are the ocean boundaries. These perhaps should not be considered interfaces but if we view the reflected field as simply an image of the incident field, then we can equate it with a direct ray that has crossed through from the other side. Therefore it sees a jump in the gradient of the sound speed that is double the gradient at the boundary. The formulas for interfaces and boundaries are essentially the same.

The other important interfaces are those associated with the oceanography. BELLHOP3D uses various piecewise approximations to the sound speed, e.g. piecewise linear in depth and in range. These lead to so-called weak discontinuities, that is, discontinuities in the derivative of the sound speed.

The curvature of these interfaces also effects the beam; however, all interfaces in BELLHOP3D are flat (locally) so we ignore these effects. The 2D version of BELLHOP does allow for a curved interpolation of the boundaries and incorporates those effects.

As beams cross such interfaces they change. Since the sound speed and density is continuous there is no change in the intensity of the beam; however, the curvature of the phase fronts of the beams changes. These changes are agnostic to the rotating trihedral. That is to say the the curvature change is directly related to the shape of the beam in the reflection plane. The reflection plane is formed by the tangent ray and the normal to the interface (both are contained in the reflection plane). Therefore, we must convert the representation of the beam from its coordinate system involving the rotating trihedral to the coordinate system of the reflection plane before converting it. Then after the interaction with the interface we need to convert it back to the coordinate system of the rotating trihedral.

Our first step is to get a new coordinate system in the reflection (or transmission plane). This is defined by the ray tangent, together with two normal vectors to the ray. The first of these normal vectors must be in the reflection plane and the second is normal to the reflection plane. If we form a normalized cross product of the ray tangent and the normal to the boundary we obtain a unit normal vector to the reflection plane:

$$
\vec{n}_{2}=\vec{n}_{\text {ReflPlane }}=\frac{-\vec{t}_{\text {ray }} \times \vec{n}_{\text {bdry }}}{\left\|\vec{t}_{\text {ray }} \times \vec{n}_{b d r y}\right\|}
$$

Taking the cross-product of the ray tangent and this normal to the reflection plane yields the other normal to the ray that is in the reflection plane:

$$
\vec{n}_{1}=\vec{t}_{r a y} \times \vec{n}_{\text {ReflPlane }}
$$

Then to express the p's into the new coordinate system we transform as follows:

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$$
\left[\begin{array}{c}
\hat{p} \\
\tilde{p}
\end{array}\right]^{\prime}=\left[\begin{array}{ll}
\vec{e}_{1} \cdot \vec{n}_{1} & \vec{e}_{1} \cdot \vec{n}_{2} \\
\vec{e}_{2} \cdot \vec{n}_{1} & \vec{e}_{2} \cdot \vec{n}_{2}
\end{array}\right]\left[\begin{array}{c}
\hat{p} \\
\tilde{p}
\end{array}\right]
$$

where the prime denotes here the values of $p$ in the new coordinate system. The same transformation is applied to the q's.

Now that we have the representation of the p's and q's in the new coordinate system we apply the curvature change formulas from Popov and Psencik (1978). (See also Cerveny and Psencik (1979); however, there is a typo in the jump equations.) These are:

$$
\begin{aligned}
{[\tilde{p}] } & =-\tilde{q} R_{1}-\hat{q} R_{2} \\
{[\hat{p}] } & =\tilde{q} R_{2}
\end{aligned}
$$

where,

$$
\begin{aligned}
& R_{1}=\frac{2}{c^{2}} \tan \alpha\left[\frac{c_{1}}{c}\right]-\frac{\tan ^{2} \alpha}{c^{2}}\left[c_{s}\right] \\
& R_{2}=\frac{\tan \alpha}{c}\left[\frac{c_{2}}{c}\right]
\end{aligned}
$$

and alpha is the angle of incidence. In addition $\left[c_{1}\right],\left[c_{2}\right],\left[c_{s}\right]$ refers to the jump in the derivatives of the sound speed across the interface (s for the derivative along the ray, 1, 2, for the derivatives in the two normal directions. The final step is to convert the ps and qs back into the raycentered coordinate system.

These formulas are also applied for boundary reflection; however, as noted above the jumps in the derivatives of the sound speeds are doubled and the sign of alpha is flipped.

Getting the signs right in all these equations is difficult because the signs of normal vectors to the boundaries and the rays need to be carefully defined.

## III. Running BELLHOP3D

A. Environmental Information: Basic input file

Our template for BELLHOP3D is the existing 2D code (BELLHOP) which is a mature code capable of handling all the practical inputs (2D sound speed profile, range-dependent bathymetry, source radiation patterns, volume attenuation, boundary reflection coefficients, etc.). As this work continues, all of these capabilities will be incorporated in BELLHOP3D. During this period, a major rewrite was done on BELLHOP3D to make the algorithm have a parallel structure to the 2D code. In the process, BELLHOP3D was set up to read essentially the same input file as BELLHOP, apart from two additional inputs required to specify the bearing lines for receivers, as well as the bearing lines used to trace the fan of beams. (These are independent; in a typical run many more beams will be traced in the azimuthal direction than there are bearing lines for the receivers.)

An example of the BELLHOP3D input file is shown below:

```
'3D Munk profile' ! TITLE
50.0 ! FREQ (Hz)
1
'CVW'
510.020000.0
    0.0 1548.52 /
    200.0 1530.29 /
    250.0 1526.69 /
    400.0 1517.78 /
    600.0 1509.49 /
    800.0 1504.30 /
1000.0 1501.38/
1200.0 1500.14 /
1400.0 1500.12 /
1600.0 1501.02 /
1800.0 1502.57 /
2000.0 1504.62 /
2200.0 1507.02 /
2400.0 1509.69 /
2600.0 1512.55 /
2800.0 1515.56 /
3000.0 1518.67 /
3200.0 1521.85 /
3400.0 1525.10 /
3600.0 1528.38/
```

```
3800.0 1531.70 /
4000.0 1535.04 /
4200.0 1538.39 /
4400.0 1541.76 /
4600.0 1545.14 /
4800.0 1548.52 /
5000.0 1551.91/
20000 1551.91/
'A~' 0.0
20000.0 1600.00 0.0 1.80.8 /
1 !NSD
1000.0 / ! SD(1:NSD) (m)
1 !NRD
10005000 ! RD(1:NRD) (m)
1001 ! NR
0.0 100.0 / ! R(1:NR ) (km)
37 !Ntheta (number of bearings)
0.0 360.0 / ! bearing angles (degrees)
'CC 3' !'R/C/I/S'
200 !Nalpha
-14.66 20 / ! alpha1, 2 (degrees) Elevation/declination angle fan
144 ! Nbeta
0 360 / ! beta1, beta2 (degrees) bearine angle fan
0.0 20500.0 99.1 ! STEP (m), ZBOX (m), RBOX (km)
'MS' 0.3 100.0 ! BeamType, epmult, Rloop (km)
15 'P' ! Nlmage IBWin
```

Sample BELLHOP3D input file.

The reader interested in the full details of this input file can find them in the Acoustics Toolbox documentation, which can be downloaded from the Ocean Acoustics Library at http:// oalib.hlsresearch.com/. However, the basic structure is probably fairly obvious. There is a sound speed profile at the beginning, followed by a series of lines that give the coordinates of the receivers in terms of receiver depth (RD), receiver range (RR), and receiver bearings. Finally, on the line showing 'CC 3', the ' 3 ' is to indicate a full 3D run. BELLHOP3D can also to a 2D run by setting that to ' 2 '. This is useful for assessing the importance of the 3D effects. We shall present some results of the code.

To display the sound-speed profile, we use the Matlab function 'plotssp.m' in the Acoustics Toolbox. The result is shown in the Figure to the right.

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Figure 1: Plot of the sound speed.

We have not yet implemented the capability to read a 3D sound speed so BELLHOP3D is currently still reading this as an analytic function implemented in FORTRAN.

## B. Environmental Information: Bathymetry

The bathymetric data is read as a matrix of ocean bottom depths, sampled on a regular grid and stored in a bathymetry file with extension '.bty'. An example bathymetry file is shown below:

```
'R'
6
-100.000000 100.000000 /
6
-100.000000 100.000000 /
4976.608187 4968.503937 4959.595960 4954.0229894956.0439564963.963964
4968.503937 4951.8072294927.272727 4906.976744 4914.8936174940.298507
4959.595960 4927.272727 4851.851852 4733.333333 4789.473684 4897.435897
4954.022989 4906.976744 4733.333333 3666.666667 4428.5714294851.851852
4956.043956 4914.893617 4789.473684 4428.571429 4636.3636364870.967742
4963.963964 4940.298507 4897.435897 4851.851852 4870.967742 4921.568627
```

Sample bathymetry file.

The first line with the letter ' $R$ ' indicates the file contains a Regular grid. (Other options are ' L ' and ' C ' which are used in the 2D BELLHOP for piecewise linear or curvilinear interpolation of bathymetry on a single slice.) The next 2 lines specify the number of points in $x$, the minimum and maximum values of $x$. The 2 lines after that have the corresponding information in $y$. Finally, we have the matrix containing the depth in meters of the bottom.

The rectilinear grid does not have to be uniform. It there is an irregular spacing then one must simply supply the full set of $x$ and $y$ coordinates. The ' $/$ ' in the above example is a standard Fortran convention that terminates the input line. When the ' $/$ ' is detected BELLHOP3D fills in a complete set of $x-y$ points using a regular spacing and treating the input values as a min and max.

We have also implemented exactly the same capability for surface elevation, allowing an 'altimetry' file to be specified. This can be used to study the effects of ocean surface waves.

The above file describes a hill. For our model run, we actually chose to sample it more finally with 41 points in both x and y . We also implemented a Matlab function called 'plotbty3D' to display it. The result is shown in the figure below.


Figure 2: Plot of the bathymetry.

In this bathymetry plot you can see how the ocean bottom is described by a pattern of rectangles. The process internal to BELLHOP3D involved quite a bit of work. In fact each rectangle is really treated as a pair of triangles. The reason for this is that three points define a plane; 4 points is too many. Thus each set of 4 points forms two triangular facets. The ray is then reflected off the triangular facet using the normal vector to that facet, which is pre-computed before the rays are traced. BELLHOP3D first identifies the particular triangle which is below the source location. The rays are then traced by stepping in arc length using a default step size. However, as the rays are traced BELLHOP checks to see if a ray will cross into the zone above another triangle. It also checks to see if the ray would cross the bottom (or surface) boundary. Then it reduces the step to land precisely on that boundary. These details are important in obtaining a precise ray trace where all the reflections are treated properly. The 2D version of BELLHOP does a similar thing; however,

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the process is obviously much simple for a single bathymetry slice versus the 2D bathymetry that BELLHOP3D processes.

## C. Environmental Information: Oceanography

It was decided to use hexahedral interpolation of the 3D oceanographic field ('brick' elements in the terminology of the finite element literature). Furthermore we assume rectilinear elements, i.e. the corner angles of the hexahedron are each 90 degrees. Then as a ray traverses through the oceanographic field, the sound speed is interpolated by 'trilinear' interpolation. To summarize this process, consider a particular ray location $\mathbf{x}$. A search is done to identify the particular segments iSegx, iSegy, iSegz, in $x, y$, and $z$, that define the element or cell containing that ray coordinate. We then construct the sound speed at the corners of a lat/lon box by interpolating the sound speed matrix, cMat, in depth:

```
c11 = cMat( iSegx, iSegy ,Layer ) + (x(3 )-zSeg(Layer ) )*
    czMat(iSegx, iSegy ,Layer)
c21 = cMat( iSegx + 1, iSegy , Layer ) + (x(3)-zSeg(Layer ) )*
    czMat(iSegx + 1, iSegy ,Layer )
c12 = cMat( iSegx, iSegy + 1, Layer ) + ( x ( 3 ) - zSeg( Layer ) )*
    czMat( iSegx, iSegy + 1, Layer )
c22 = cMat( iSegx + 1, iSegy + 1, Layer ) + ( x( 3 ) - zSeg( Layer ) ) *
    czMat( iSegx + 1, iSegy + 1, Layer )
```

where czMat is the gradient in depth of the sound speed, which is calculated when the matrix of sound speed values is first read in.

Next we calculate the proportional distances of the ray coordinate within the active cell. Denoting the proportional distances in x and y by s1 and s2 respectively, we calculate:

```
s1 = (x(1) - xSeg( iSegx ) ) / ( xSeg(iSegx + 1 ) - xSeg( iSegx ) )
s2 = (x(2) - ySeg(iSegy ))/(ySeg(iSegy + 1 ) - ySeg(iSegy ) )
```

where $x$ Seg and ySeg are the coordinates of the segment defining the active cell. We then interpolate the soundspeed in the y direction, at the two endpoints in x (top and bottom sides of rectangle):

$$
\begin{aligned}
& \mathrm{c} 1=(1.0 \mathrm{DO}-\mathrm{s} 2)^{*} \mathrm{c} 11+\mathrm{s} 2{ }^{*} \mathrm{c} 12 \\
& \mathrm{c} 2=(1.0 \mathrm{D} 0-\mathrm{s} 2)^{*} \mathrm{c} 21+\mathrm{s} 2{ }^{*} \mathrm{c} 22
\end{aligned}
$$

Finally we interpolate the soundspeed in the $x$ direction, at the two endpoints in $y$ (top and bottom sides of the rectangle)

$$
c=(1.0 D 0-s 1)^{*} c 1+s 1^{*} c 2
$$

The details here are not terribly important but hopefully this outline conveys the notion that the soundspeed is interpolated linearly within a cell. Gradients are also easily calculated in this process.

An example BELLHOP3D input file (KoreanSea3d_ray.env) to invoke an oceanography field is shown below. The key parameter is ' H ' in the SSOPT line, which indicates hexahedral interpolation. The lines with depths from 0 to 5500 would normally contain the actual sound speed profile; however, they are ignored when the ' H ' option is selected.
'Korean Sea 3D' ! TITLE
250.0

1
'HVW'
510.05500 .0

0 /
10 /
20 /
30 /
50 /
75 /
100 /
125 /
150 /
200 /
250 /
300 /
400 /
500 /
600 /
700 /
800 /
900 /
1000 /
1100 /
1200 /
1300 /
1400 /
1500 /
1750 /
2000 /
2500 /
3000 /
3500 /
4000 /
4500 /
! FREQ (Hz)
! NMEDIA
! SSPOPT (Analytic or C-linear interpolation)
! DEPTH of bottom (m)

5000 /
5500 /
'A~' 0.0
5500.01600 .000 .01 .80 .8 /

1 ! NSD
100.0 / ! SD(1:NSD) (m)

1 ! NRD
25.0 ! $\quad$ RD(1:NRD) (m)

1001 ! NR
0.050 .0 ! $\mathrm{R}(1: \mathrm{NR})(\mathrm{km})$

6 ! Ntheta (number of bearings)
0.0360 .0 / ! bearing angles (degrees)
'RC 3' !'R/C/I/S'
25 ! Nalpha
-10 10 / ! alpha1, 2 (degrees) Elevation/declination angle fan 4 ! Nbeta
90270 / ! beta1, beta2 (degrees) bearine angle fan
0.05500 .0 50. ! STEP (m), ZBOX (m), RBOX (km)
'MS' 0.3100 .0 ! BeamType, epmult, Rloop (km)
15 'P' ! Nlmage IBWin

Sample BELLHOP3D input file (KoreanSea3d_ray.env) when 3D oceanography is included.

The actual sound speed field is then provided as an additional input file with the extension 'SSP' (in this case 'KoreanSea3d_ray.ssp'). The format of the SSP file is as follows:


These lines are then followed with the matrix of sound speed values. In this cases we consider a 3D volume consisting of 151 points in $x, 169$ points in $y$, and 33 points in depth. To avoid wasting space, we deleted some of the data as indicated by the ellipses.

Besides the hexahedral option, a full set of SSP interpolation routines was also implemented for purely stratified problems. Thus BELLHOP3D can also be easily run with a single SSP, which can be interpolated in several ways (c-linear, n2-linear, cubic spline) in depth, just like the 2D version of BELLHOP.
D. Ray trace run: Nx2D

Here we are following a typical procedure where we first plot the environmental information to make sure that the problem is set up as intended. Next we usually like to look at a ray trace to see how the sound energy is traveling. This is done by selecting the RunType to be ' $R$ ' in BELLHOP3D. For our initial run, we also select the 2D option in BELLHOP3D. After the run is completed, BELLHOP3D echoes information to a print file as shown below.

```
BELLHOP- 3D Munk profile
    frequency = 50.00 Hz
    Dummy parameter NMedia =
        C-linear approximation to SSP
        Attenuation units: dB/wavelength
        VACUUM
    Depth = 20000.000000000000 m
    C-Linear SSP option
    Sound speed profile:
            0.00 1548.52
        200.00 1530.29
        250.00 1526.69
        400.00 1517.78
        600.00 1509.49
        800.00 1504.30
        1000.00 1501.38
        1200.00 1500.14
        1400.00 1500.12
        1600.00 1501.02
        1800.00 1502.57
        2000.00 1504.62
        2200.00 1507.02
        2400.00 1509.69
        2600.00 1512.55
        2800.00 1515.56
        3000.00 1518.67
        3200.00 1521.85
        3400.00 1525.10
        3600.00 1528.38
        3800.00 1531.70
        4000.00 1535.04
        4200.00 1538.39
        4400.00 1541.76
        4600.00 1545.14
        4800.00 1548.52
```

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| Number of source depths = Source depths (m) $1000.00$ |  |  |  |
| :---: | :---: | :---: | :---: |
| Number of receiver depths $=$ Receiver depths (m) $1000.00$ |  |  |  |
| Number of ranges = | 1001 |  |  |
| Receiver ranges (km) |  |  |  |
| 0.000000 .100000 | 0.200000 | 0.300000 | 0.400000 |
| 0.5000000 .600000 | 0.700000 | 0.800000 | 0.900000 |
| 1.000001 .10000 | 1.20000 | 1.30000 | 1.40000 |
| $1.50000 \quad 1.60000$ | 1.70000 | 1.80000 | 1.90000 |
| 2.00000 |  |  |  |
| .. 100.000000 |  |  |  |
| Number of bearings = | 5 |  |  |
| Receiver bearings (degrees) |  |  |  |
| 0.0000075 .0000 | 150.000 | 225.000 | 300.000 |

```
Ray trace run
Cartesian beams
Point source (cylindrical coordinates)
Rectilinear receiver grid: Receivers at rr( : ) x rd( : )
N x 2D calculation (neglects horizontal refraction)
Number of beams in elevation = 30
Beam take-off angles (degrees)
\begin{tabular}{rrrrr}
-14.6600 & -13.4648 & -12.2697 & -11.0745 & -9.87931 \\
-8.68414 & -7.48897 & -6.29379 & -5.09862 & -3.90345 \\
-2.70828 & -1.51310 & -0.317931 & 0.877241 & 2.07241 \\
3.26759 & 4.46276 & 5.65793 & 6.85310 & 8.04828
\end{tabular}
        9.24345
Replacing beam take-off angles, beta, with receiver bearing lines, theta
Number of beams in bearing = 5
Beam take-off angles (degrees)
    0.00000 75.0000 150.000 225.000 300.000
```

```
Step length, deltas = 2000.0000000000000 m
Maximum ray Depth, Box%z = 20500.000000000000 m
Maximum ray range, Box%r = 99100.000000000000 m
Type of beam = M
Standard curvature condition
Epsilon multiplier 0.29999999999999999
Range for choosing beam width 100.000000000000000
Number of images, Nimage = 1
Beam windowing parameter = 5
Component = P
*********************************
Using bottom-bathymetry file
Regular grid for a 3D run
Number of bathymetry points in x-direction 41
    xMin (km) = -100.000000000000000 xMax ( km ) =
100.000000000000000
Number of bathymetry points in y-direction 41
    yMin (km) = -100.000000000000000 yMax ( km ) =
100.000000000000000
Minimum width beams
HalfWidth1 = 977.65495322067386
HalfWidth2 = 977.65495322067386
epsilonOPT1 = ( 0.0000000000000000 , 150138159.23471510 )
epsilonOPT2 = ( 0.0000000000000000 , 150138159.23471510 )
EpsMult = 0.29999999999999999
ibeta 1 0.0000000000000000
ibeta 2 75.000000000000000
ibeta 3 150.00000000000000
ibeta 4 225.00000000000000
ibeta 5 300.00000000000000
CPU Time = 0.893E-01s

Following the pattern of all the models in the Acoustics Toolbox, the user's input data is echoed to the print file as soon as it is read. Thus, if there is an error, one can see at precisely which line of the input file the error occurred. In addition, we have structured BELLHOP3D to use the same subroutine as the standard BELLHOP to read in this data. That routine is called with a flag to indicate whether the input file has the additional 3D information needed for a BELLHOP3D run.

Notice the CPU time displayed on the last time, which indicates that this run took less than a tenth of a second on a MacBook Pro laptop computer.

The key output of this run is a ray file with extension '.ray'. The format differs slightly from a 2D BELLHOP ray file in that the rays are now trajectories in \(x-y-z\) space, rather than \(r-z\) (range-depth) space. We have also implemented a Matlab function called 'plotray3D' to display the rays. The resulting display is shown in Figure 6.

Like the 2D BELLHOP, 'plotray3D’ automatically colors the ray based on whether it hits the top and/or bottom boundaries. This makes for a clearer display of where the energy is traveling. The source depth here is 800 meters, which is close to the sound channel axis. This leads to an interesting pattern of refracted rays within the water column, showing the well-known convergence zone effect. The results cross-check with the 2D version of BELLHOP, indicating that BELLHOP3D is implemented correctly.

If we rotate this ray plot to look straight down on the ocean, we get the display shown below. Note that all the rays lie precisely in vertical planes defined by the bearing lines. This is because we ran BELLHOP3D in its 2D mode, so horizontal refraction is neglected. Finally, we can display both the rays and the bathymetry in a single plot as shown in Figure 7. Here you can see more clearly how the rays are being reflected off the ocean bottom.

Figure 6: Ray trace in the Nx2D mode (upper plot is a side view, lower plot is a top view).


Figure 7: Ray trace plotted over the bottom bathymetry.

\section*{E. Ray trace run: Full 3D Mode}

Next we illustrate BELLHOP3D in its full 3D mode. As mentioned above, this is accomplished by changing a single character in the input file. The print file is essentially the same so we will not repeat it here. However, the 3D option is a bit slower, requiring 0.17 seconds. The ray plot is shown below.

Again, to see more clearly the horizontal refraction effects we switch to a view from above where now you can see the rays have refracted out of the vertical plane due to multiple reflections from the hill, as well as the gradients in the ocean sound speed. The paths going over the hill naturally show more out of plane refraction.

Figure 8: Ray trace in the 3D mode (upper plot is a side view, lower plot is a top view).

\section*{F. Transmission Loss}


To do a transmission loss calculation we change the RunType from 'R' (ray trace) to 'C' (coherent transmission loss). In addition, we increase the number of takeoff angles in the vertical from 30 to 200. For a ray trace we typically find the plots are too cluttered if we use this many rays; however, the TL results are more accurate with finer ray fans.

Similarly, to get a well sampled transmission loss plot in the lat./long. plane, we increase the number of bearing lines for the receivers to be every 10 degrees. For a 2D run of BELLHOP 3D, the code automatically uses this same set of angles for the azimuthal launch directions of the rays. However, for a 3D run we increase this to 144 to get finer sampling. This is necessary because the beams are now modeling the horizontal refraction. The resulting 2D and 3D plots are shown in Figure 9. The run time for the Nx2D calculation was 1.5 seconds versus 26.2 seconds for the full 3D calculation. The 3D calculations are slower because the governing equations are more complicated, but also because a finer sampling of rays is needed in the azimuthal direction. Currently we are not taking advantage of the multiprocessors that are available.

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Figure 9: BELLHOP3D transmission loss (upper plot is an Nx2D calculation, lower plot is a 3D calculation).

\section*{IV. Test Cases}

A number of test cases have been used to benchmark BELLHOP3D and also demonstrate its capabilities. These test cases are provided in their entirety in the directory at/tests/bellhop3d. On overview of the test cases follows:
- Free space: This simple point source problem is used to verify that the geometric beams are space-filling in all directions and have the correct angle-dependent scale factors for transmission loss.
- Perfect Wedge: This case approximates a continental shelf region as a wedge. The term 'perfect' refers to the fact that the boundaries are perfectly reflecting.
- Truncated Wedge: A case taken from the text by Jensen, Kuperman, Porter, and Schmidt; similar to the perfect wedge treated in the last report but has a penetrable bottom.
- Seamount: Another benchmark case with an analytic solution.
- Harvard Eddy Scenario: A test case originally used for the FOR3D parabolic equation model.
- Rotated Munk Profile: A classical deep-water case turned on its side. Essentially exact solutions are available from other numerical models; however, turning it on its side exercises all the out-of-plane effects in BELLHOP3D.
- Rotated Lloyd's Mirror: An analytic solution exists. Again, the rotation allows us to test the treatment of out-of-plane effects and the complicated effects of beam reflection off sloped boundaries.

In the following sections we provide more detail on these benchmark cases.

\section*{A. Free-Space and Half-Space Propagation}

An isovelocity problem is the simplest test case. However, it is not entirely trivial. The simple spherical decay of the field allows any small flaws in the beam widths to show up. If the beams are either too narrow or too wide, the field will show obvious irregularities. In Fig. 10 below we present in succession the results of a wavenumber integration code (SCOOTER), and BELLHOP3D run with Geometric hat and Geometric Gaussian beams. The agreement is excellent except in the very near field where we see an artifact due to a far-field approximation normally used in wavenumber integration models.


Figure 10: Field due to a point source in free space. Top panel is a reference solution calculated using SCOOTER. The middle panel is calculated using BELLHOP3D with Geometric Hat Beams. The Lower panel is calculated using BELLHOP3D with Geometric Gaussian beams.

Figure 11 shows the BELLHOP3D results taken for a slice at fixed depth; i.e., a plan view. In the upper panel we have used the hat-shaped beams. In the lower panel we have used Gaussian beams. Collectively these results demonstrate that the code is producing a beautiful image of the field due to a point source regardless of the slicing plane and with both hat and Gaussian beams. This might seem like an easy test case but in fact it is a fairly stringent test of the model function.


Figure 11: Acoustic field calculated by BELLHOP3D at a fixed receiver depth. The upper panel uses Geometric Hat beams while the lower panel uses Geometric Gaussian beams.

Next we introduce upper and lower halfspaces with a sound speed of \(1750 \mathrm{~m} / \mathrm{s}\) over an ocean with a sound speed of \(1500 \mathrm{~m} / \mathrm{s}\). There is no density contrast. The result in Fig. 3 shows the expected interference pattern. The left panels are a reference solution from the SCOOTER wavenumber integration code. The right panels are the results of the new BELLHOP3D code. The upper panels are for a lower halfspace the the lower panels are for an upper halfspace.

One can see a variation in the energy vs. angle that is due to the corresponding variation in the reflection coefficient for a halfspace. The precise positions of the lobes in the resulting interference pattern are extremely sensitive to the phase. The excellent agreement between all the results gives a lot of confidence in the treatment of the boundary reflection.

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Figure 12: Comparison of BELLHOP3D to a reference solution calculated by SCOOTER for lower and upper halfspaces

\section*{B. The Perfect Wedge}

The perfect wedge is a classic problem has been studied by various people (Bradley and Hudimac, Buckingham, Tolstoy, ...). Here we use an analytic formula reported by Buckingham, which we (Xavier Zabal at HLS) coded up in Matlab. We consider a case studied previously by Tolstoy, Buckingham, and Doolite with a 10 Hz source at 19.1 km from the apex. The source depth is 8 m and the receiver depth is 80 m (Fig. 13). The water depth at the source location is 400 m and the pressure vanishes at both the top and bottom boundaries. The resulting transmission loss is shown in Fig. 14.


Figure 13: Schematic of the wedge problem.

Figure 14: Transmission loss for the wedge problem derived from the analytic solution.

To interpret this figure it is useful to start with the BELLHOP3D ray traces. Here we have selected 10 rays in the azimuthal direction going from -90 to +90 degrees. In the vertical we chose 21 beams from -20 to +20 degrees. The top view of the rays is shown in Fig. 15 and the perspective view is shown in Fig. 16. Here we can see that the energy that propagates upslope is gradually refracted back downslope through repeated interactions with the sloping seafloor. The analytic

formula of Buckingham (implemented by Xavier Zabal at HLS) shows that the field is actually formed from 5 sets of wedge modes. The major bands of energy in the TL plot correspond to the caustics associated with the ray paths of the modes.

Finally, in Fig. 17 we show the TL as calculated by BELLHOP3D. The result shows nicely the general characteristics of the field; however, it is not reproducing the fine details of the interference pattern. This indicates that there is still some error in the code that needs to be fixed. It should be noted that this is an extraordinarily difficult test case. We are going out a very long distance and generating a large number of bottom reflections. The vertical angles of propagation go to 90 degrees and, because the boundaries are perfectly reflecting, all the upslope energy is returned downslope where it forms a delicate interference pattern with itself.


Figure 15: Top view of the rays for the wedge problem.


Figure 16: Perspective view of the rays for the wedge problem.



Figure 17: Transmission loss for the wedge calculated by BELLHOP3D; hat-beams (upper) and Gaussian beams (lower).

\section*{C. Truncated Wedge}

This test case explores the effects of a sloping bottom on acoustic propagation. The geometry considered is that of a truncated wedge as shown in the Fig. 18.


Figure 18: Schematic of the truncated wedge problem (from Jensen, Kuperman, Porter, and Schmidt).

The angle of the wedge is 2.86 degrees and runs from a shallow depth of 20 m to a depth of 380 m in a range of 7.2 km . The sound speed in the water is \(1500 \mathrm{~m} / \mathrm{s}\) while the homogeneous fluid bottom has a speed of \(1700 \mathrm{~m} / \mathrm{s}\), a density of \(1500 \mathrm{~kg} / \mathrm{m}^{3}\), and an attenuation of \(0.5 \mathrm{~dB} / \lambda\). The source transmits at a frequency of 25 Hz , and is located halfway down the slope at a depth of 40 m . At the source position the depth of the water is 200 m . The acoustic field is sampled at a depth of 30 m around the source.

The BELLHOP3D environment file for this case is shown below.
```

'TWedge profile, 3D Hat beams 721 x 721' ! TITLE
25.0 ! FREQ (Hz)
1 ! NMEDIA
'CVW' ! SSPOPT (Analytic or C-linear interpolation)
0 0.0 20000.0 ! DEPTH of bottom (m)
0.0 1500.00 /
20000 1500.00 /
'A~' 0.0

```

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```

20000 1700.0 0.0 1.5 0.5 /
1 ! Nsx number of source coordinates in x
0.0 / ! x coordinate of source (km)
1 ! Nsy number of source coordinates in y
0 ! ! y coordinate of source (km)
1 ! NSD
40.0 / ! SD(1:NSD) (m)
1 ! NRD
30 / ! RD(1:NRD) (m)
1001 ! NR
0.0 25.0 / ! R(1:NR ) (km)
1 4 4 1 ~ ! ~ N t h e t a ~ ( n u m b e r ~ o f ~ b e a r i n g s )
-90 90 / ! bearing angles (degrees)
'CG 3' ! 'R/C/I/S'
1441 ! Nalpha
-89 89 / ! alpha1, 2 (degrees) Elevation/declination angle fan
1441 %144 ! Nbeta
-90 90 / ! beta1, beta2 (degrees) bearine angle fan
0 25 4 20500.0 ! STEP (m), Box%x (km) Box%y (km) Box%z (m)

```

For this test case there is no analytic solution for the propagation, but the BELLHOP3D results are compared to the Nx2-D and 3-D solutions reported by Sturm¹, using a 3-D PE model.

Figure 19 presents the results obtained by Sturm for the propagation loss computed using an Nx2D method in the top panel, and with a full 3D PE method in the bottom panel.

\footnotetext{
\({ }^{1}\) F. Sturm, "Numerical study of broadband sound pulse propagation in three dimensional oceanic waveguides", J. Acoust. Soc. Am. 1117, 1058-1079 (2005)
}

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Figure 19: Comparison of Nx2D and 3D results from the Sturm PE for the truncated wedge (figure from Jensen, Kuperman, Porter, and Schmidt).


Figure 20: Comparison of Nx2D and 3D results from BELLHOP3D for the truncated wedge.

These results may be compared to the BELLHOP3D outputs for the Nx2-D (top) and full 3D (bottom) methods shown in Fig. 20.

Both results show a significant difference between the Nx2D and 3D methods, and the horizontal refraction effects are clearly visible and qualitatively similar in the two models. If we look closely at the interference pattern we see excellent agreement towards the right side (cross-slope); however, in the lower left (down-slope) there are some clear differences. There are approximations in both BELLHOP3D and in Sturm's PE (narrow-angle) so from this test we cannot be sure which is the more accurate solution. However, there is a significant difference in the starting fields for the two
models. Sturm used a starting field consisting of just the first few modes; however, in BELLHOP3D we use a point source starting field.

To try to resolve this discrepancy, the propagation down the slope was compared to the prediction of the RAM acoustic model. Figure 21 shows the output of the RAM model next to the BELLHOP results. The agreement is excellent, but detailed analysis shows that there is a small rangedependent shift between both models, a difference due to the fact that the beam-shift effect caused by the penetrable bottom is not implemented in BELLHOP3D right now. There is also a bit of noise in the BELLHOP3D results near the boundary which is a normal characteristic that can be reduced by using additional beams.


Figure 21: Comparison of RAM (PE) and BELLHOP3D for a downslope track in the truncated wedge.

Another model we can compare to is Ding Lee's FOR3D code. The result from FOR3D for this case is shown in Fig. 22. There is an artifact in the lower right corner whose cause is not known at the moment. However, the rest of the figure is in general agreement with BELLHOP3D. Here we can also see excellent agreement in the lower-left or downslope region.

Figure 22: Results from FOR3D (3D PE model) for the truncated wedge.

In summary, we have presented results from 4 different acoustic models (BELLHOP3D, RAM, Sturm PE, FOR3D). The results generally agree; however no single pair shows really the level of agreement we would hope for. Given that BELLHOP3D agrees with RAM and FOR3D in the downslope region and with the Sturm PE in the cross-slope region, we conjecture that BELLHOP3D is providing the best solution throughout the domain. The two PEs use different

operator splittings and the differences between them are probably largely due to that. However, the difference in the starting field is another important issue.

\section*{D. Seamount}

This test case involves propagation around a conical seamount that extends all the way to the sea surface. Both the sea surface and the seamount are assumed to be perfect reflectors, and this assumption permits an analytical solution to the Helmholtz equation based on a separation of variables technique, as developed by Buckingham².

For this test it is assumed that the seamount slope gradient is 75 degrees (channel angle of 15 degrees). The source is at a depth of 100 m and located at a range of 3 km to the right of the seamount apex, and the transmitted frequency is 20 Hz . The acoustic field around the seamount is also sampled at a depth of 100 m .

The BELLHOP3D environment file used to run this problem is shown below:
```

'Seamount3D, Gaussian 361x601 rcvrs, 361x301 beams' ! TITLE
5.0 ! FREQ (Hz)
1 ! NMEDIA
'CVW' ! SSPOPT (Analytic or C-linear interpolation)
0 0.0 20000.0 ! DEPTH of bottom (m)
0.0 1500.00 /
20000 1500.00 /
'R~' 0.0
1 ! Nsx number of source coordinates in x
3.0 / ! x coordinate of source (km)
1 ! Nsy number of source coordinates in y
0.01 / ! y coordinate of source (km)
1 ! NSD
100.0 / ! SD(1:NSD) (m)
1 ! NRD
100 / ! RD(1:NRD) (m)
6 0 1 ~ ! ~ N R
0.0 12.0 / ! R(1:NR ) (km)
3 6 1 ~ ! ~ N t h e t a ~ ( n u m b e r ~ o f ~ b e a r i n g s )
90 180 / ! bearing angles (degrees)
'CB 3' ! 'R/C/I/S'
301 ! Nalpha
-89 89/ ! alpha1, 2 (degrees) Elevation/declination angle fan
361 ! Nbeta
90 180 / ! beta1, beta2 (degrees) bearing angle fan
100 9 9 20500.0 ! STEP (m), Box%x Box%y Box%z (m)

```

\footnotetext{
\({ }^{2}\) M.J. Buckingham, "Theory of acoustic propagation around a conical seamount", J. Acoust. Soc. Am. 80, 265-277 (1986).
}

Figure 23 presents a top view of the trajectory of 15 rays leaving the source at angles between 135 and 225 degrees relative to the horizontal axis. The left panel corresponds to rays with a grazing angle of +20 degrees (downward from the source), while the right panel corresponds to rays with an upward grazing angle of 8 degrees.

Figure 24 shows the results of the BELLHOP3D run along with the results of the analytical model. In the BELLHOP3D case the coordinate system is centered at the source position while for the analytical model the coordinates are centered at the apex of the seamount. The overall energy levels show a reasonable match; however, the BELLHOP3D result is more 'broken up'. We assume this is the 'disco ball' effect, i.e. due to the fact that BELLHOP3D uses a facetted description of the bottom rather than the perfect cone used in the analytic formulation. Real oceans are or course neither perfect cones nor perfectly facetted.


Figure 23: Top view of the BELLHOP3D rays for the seamount problem.

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Figure 24: Comparison of BELLHOP3D transmission loss to the analytic model.

\section*{E. The Harvard Case Eddy Scenario}

Another approach to validating BELLHOP3D is to compare it to other numerical solutions. However, there are not so many models capable of handling 3D scenarios. Generally, this remains a research issue even though basic 3D models were developed back in the 1980s. The reason such models have not really taken off is 1) environmental information (oceanography especially) has not readily been available, 2) run times have been excessive for 3D models. BELLHOP3D is very practical for such 3D problems largely because of the basic efficiency of ray-based models. Meanwhile, the speed of typical computers has increased perhaps by a factor of 10,000 since those original models were developed. Nevertheless, it remains difficult to run other model types at reasonable frequencies to generate comparison solutions. Generally, we are forced to take artificially low frequencies (where BELLHOP3D tends to be less accurate) to create a test problem that can be solved by both BELLHOP3D and another model type.

We have chosen to use Ding Lee's FOR3D Parabolic Equation (PE) model here. This is one of the few 3D PE models available and is openly distributed on the Ocean Acoustics Library. However, it is more that 30 years old and there was considerable work required to install it. While the code was generally carefully implemented, there were a number of problems that showed up with the code, e.g. DO loops using real variables and issues with uninitialized variables. In the process of installing this, she also converted it to a more modern FORTRAN 95 version. We present here our first test case of the model. BELLHOP3D comparisons have not been made yet.

The scenario is the 'Harvard' Case provided with the code, which involves long-range propagation (over about 4 degrees in latitude and longitude) in an area of the North Atlantic where the Gulf Stream is present, spinning off numerous warm and cold core eddies. The actual site with a cross showing the source location is shown in Fig. 25. There is a mild bathymetric variation; however, the key feature of interest is probably the oceanography with the various eddies. The SSP on a horizontal slice is shown in Fig. 26; the SSP on a vertical slice at the source location is shown in Fig. 27. Finally Figs. 28 and 29 show the transmission loss computed by FOR3D on both vertical and horizontal slices.

These are just preliminary results that show that we have been able to install the FOR3D model successfully.

Figure 25: Model site for the Harvard scenario.

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Sound Speed Contours


Figure 26: Sound speed on a horizontal slice.

Figure 27: Sound speed profile at the source location for the Harvard scenario.

Figure 28: Transmission loss on a vertical slice for the Harvard scenario.

Figure 29: Transmission loss on a horizontal slice for the Harvard scenario.



\section*{F. Rotated Munk Profile}

The Munk profile is a classic deep water profile as shown in Fig. 30 below. We have used this frequently as a test and have reference solutions calculated by a wide variety of models. In Fig. 31 we show the transmission loss as calculated by BELLHOP3D using the Geometric Hat-Beam option. The important test on this figure was to verify that it reproduced exactly the result from the standard 2D BELLHOP. In fact, both 2D and 3D versions of BELLHOP call the same influence routine and the results agree at the pixel level. (See also the results presented in Progress Report 5.)

However, the main motivation for this case was to verify that BELLHOP3D was handling all the lateral interfaces in the oceanography correctly. For this to work properly we also had to add logic to BELLHOP3D to detect all the oceanography interfaces, not just those in depth. (As noted previously, the ray traces do not come out precisely if the rays are allowed to step over interfaces.) In addition, the logic for beam curvature changes had to be implemented. This involves rotating the beam equations into a coordinate system relative to the plane of the incident ray on the interface. The point of this discussion is just to emphasize that there is quite a bit of code required to handle all of this properly.

To capture these effects, we simply rotate the sound speed profile so that the gradients are all in the \(x\)-direction, not the usual z-direction. The profile is completely unphysical but that does not matter for this benchmark objective. The resulting transmission loss from BELLHOP3D is shown in

Fig. 32. Note that it is virtually identical to the previous figure except for the axis rotation. One difference that can be seen is that the new result appears fuzzier at long ranges. This is because BELLHOP3D uses a polar coordinate system with radials emanating from the source- at longer ranges the radials spread further apart.

The excellent match here may seem anticlimactic; however, initial results showed a variety of problems and a great deal of work was necessary to reach this exact agreement.


Figure 30: Plot of the Munk sound speed profile.

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BELLHOP3D- Munk \({ }_{3}\) Dz 3D


Figure 31: Transmission loss for the Munk profile calculated by BELLHOP3D.


Figure 32: Transmission loss for the Rotated Munk Profile calculated by BELLHOP3D.

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\section*{G. Korean Seas}

Here we consider a modeling exercise for the Korean Seas (East, South, West). An oceanographic field was derived from the HYCOM (Hybrid Coordinate Ocean Model). The HYCOM model is discussed on http://hycom.org/. The associated consortium provides global oceanographic forecasts on a daily basis. For this example, we downloaded the global forecast for Julian day 204 of 2011 with output provided at about 12.5 points per degree. A short Matlab script was written to 1) read the global NetCDF data for salinity and temperature, 2) extract a desired operational area, 3) convert the salinity and temperature to sound speed, and 4) write the SSP field in the above described format that BELLHOP3D requires. The resulting field is plotted in Figure 33 below.

BELLHOP3D requires the SSP field to be provided on a Cartesian grid ( \(x-y-z\) ) coordinates. The lower left corner of the box is arbitrarily selected as the coordinate origin. The results on the Cartesian grid are plotted in the lower panel of Figure 33.

In Figure 34 we display also a section at constant latitude ( \(\mathrm{y}=700 \mathrm{~km}\) ) to illustrate the upward refracting nature of the profile. The 'white' sections in the plot a NaNs corresponding to areas of land.

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Figure 33: Horizontal slice (at a depth of 50 m ) of the sound speed field. Top panel presented with latitude/longitude in degrees; bottom panel presented in Cartesian coordinates.


Figure 34: Vertical slice (at a latitude of 700 km ) of the sound speed field.

The other critical piece of environmental information is the bathymetry. The actual bathymetry used in the HYCOM simulations can also be downloaded from the HYCOM website. However, rather than sort out the format of that file, we download the bathymetry data for this same box from the National Geophysical Data Center website (http://www.ngdc.noaa.gov/ ). In particular we select the ETOPO1 database which provides data on a 1' grid. The data was downloaded in the standard xyz format. A routine was then written to read the xyz file and write it in the format the BELLHOP3D requires for a bathymetry file.

Thus, we now have simple tools that enable us to download an operational area anywhere on the globe and immediately generate the SSP and bathymetry files that BELLHOP3D requires. The bathymetry for this test case is shown in Figure 35.

Figure 35: Bathymetry for the Korean Seas from the ETOPO1 database.

Figure 36: Ray trace over the bathymetry, showing the location of the source.


Figure 37: Ray trace over the bathymetry. In the lower panel faceting has been turned on to show more clearly the interaction of the rays with the bathymetry.

We selected a source location in the East Sea at (1000 km, 750 km ) in our \(x-y\) grid and traced a fan of rays as shown in Figure 36. Figure 38 shows a blowup (with surface faceting both on and off) where you can clearly see the role of the bathymetry. With the faceting turned off, the picture is less cluttered, making it easier to see the actual ray trace. However, the faceting allows one to see more clearly how the actual bathymetry causes the rays to refracts out of the vertical plane.

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Figure 38: Ray trace over the bathymetry for a source closer to shore.

To accentuate the horizontal refraction effects we experimented with different source positions. We show here another example, where the source location is shifted slightly towards the coast at (970 km, 750 km ). Again we present the results with faceting turned off and on (Fig. 38). Here we can see dramatic refractive effects in the horizontal plane. Those effects are primarily due to the bathymetry since the coarsely sampled oceanographic field does not have very strong gradients.

Note also that some of the rays are completely turned around. In practice the energy associated with these rays will be severely attenuated due to bottom losses.

Figures 39 and 40 demonstrate the multi-source capability for the ray traces. In this case 4 sources have been placed at different lat/longs and rays were traced in the North, South, East, and West directions. The upper plot was done using the Nx2D option so there is no horizontal refraction.

Figures 41 and 42 show slices of the transmission loss field using the Nx2D option, and full 3D option. The range-depth slice is taking along the 0 degree (eastward) bearing for the first source. The lat/long slice is taken at a receiver depth of 400 m .

Differences between the two plots are partially due to horizontal refraction. However, we expect that there are also some differences here due to the different beam types used. The Nx2D model uses Geometric Beams here, while the 3D run uses Cerveny beams. Typically we find the Geometric Beams are more reliable.


Figure 39: Ray trace (Nx2D option) showing the multi-source capability.


Figure 40: Ray trace (3D option) showing the multi-source capability.



Figure 41: Transmission loss (Nx2D option, side and plan views) showing the multi-source capability.


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Figure 42: Transmission loss (3D option, side and plan views) showing the multisource capability.

\section*{H. Incoherent and Semi-Coherent TL options}

As discussed in the text Computational Ocean Acoustics (Jensen, et al., Springer-Verlag), incoherent and semi-coherent options are often useful for treating certain types of sources. Briefly, the motivation for this is that particularly at higher frequencies; e.g., mid-frequency sonar, the transmission loss pattern is a complicated field that results from constructive and destructive interference. That field is extremely sensitive to the details of the oceanography and bathymetry-more detail than the oceanography and bathymetry can provide--and therefore is not really conveying meaningful information. Instead what is often desired is some kind of average results that captures the general level of the field.

As it turns out, there are also computational benefits in computing such an averaged quantity since the sampling requirements (number of beams) is greatly reduced.

Two types of averaged results have been found to be useful. The first is an incoherent field in which the phase of the rays or beams is discarded. This produces a very smooth field. The second is a semi-coherent field in which the phase is discarded except that a beam pattern is applied at the source representing the Lloyd mirror pattern that results from the interference of the direct field and the surface reflected field. Thus one captures a dominant interference effect that may be a robust feature, but ignores other effects that might not be stable with respect to small fluctuations in the environment or source/receiver geometry. The choice of which of these options is best depends on the source frequency and other features of the scenario.

We have implemented this incoherent and semi-coherent option in 1) geo-hat, 2) geo-hat in Cartesian coordinates, 3) geo-Gaussian beams. To introduce the problem we begin with the standard coherent result. Our test case is the Munk profile (standard deep water profile) and we present BELLHOP3D results in Fig. 43 below. The upper panels are for the Hat-shaped beams, while the lower panels use Gaussian beams. The left panels use the Nx2D option and the right panels use the full 3D option. We have solved this problem previously with a variety of other codes and can say that the results are excellent. One may also note the excellent agreement between the Nx2D and 3D results and the improved accuracy in the Gaussian beam results.

Figure 43: Fully coherent calculations of the TL using BELLHOP3D with hat and Gaussian beams and with Nx2D and full 3D options.

Figure 44 shows the results with the new incoherent and semi-coherent options in BELLHOP3D. We show first the results with the geometric, hat-shaped beams. Note how the incoherent result in the top panel shows a much more smoothed field that conveys the general energy levels. The lower panel shows the semi-coherent option. Note here how it reproduces an interference pattern

that is reminiscent of the fully coherent solution, but differs in some details. As discussed above, it is capturing the interference due to the surface image, but not other interference features.

Figure 44: Incoherent and semi-coherent TL fields calculated by BELLHOP3D using hat-shaped beams.

The remaining figures are perhaps less interesting but show that we have also implemented this capability for other beam types. Figure 45 shows the same pair of results but using the geo-hat beams in Cartesian coordinates. These results should be, and are, identical to the previous set. The significance of the implementation in Cartesian coordinates is simply that the computer time is significantly reduced.

Figure 45: Incoherent and semi-coherent TL fields calculated by BELLHOP3D using hat-shaped beams in Cartesian coordinates.


Finally, in Fig. 46 we show the results using Gaussian beams. As usual the Gaussian beams smooth out artifacts in the field. However, there is a level bias. This is due to the fact that all the beams are added incoherently when in fact beams of the same group, that is, with the same number of upper and lower turning points should always be added coherently. It is possible to derive a simple formula that compensates for that.

Figure 46: Incoherent and semi-coherent TL fields calculated by BELLHOP3D using Gaussian beams.


\section*{V. Summary}

BELLHOP3D is now a fully capable production model at the level of the widely used BELLHOP code that has been used previously for 2D problems. This report focuses on the aspects that are new for the 3D capability; however, the code has been structured in a manner that parallels the earlier BELLHOP code. We refer the user to the BELLHOP manual for additional info in using the code.

\section*{Acknowledgments}

BELLHOP3D was originally developed at the Naval Research Laboratory around 1985 as a preliminary research code. This new version of BELLHOP3D including the User Guide has been largely supported by the Agency for Defense Development, Republic of Korea. It greatly extends the original by including options for gridded environmental data supplied by the user. It also provides a number of new outputs such as eigenrays and 'arrivals files'. Finally additional beam types were implemented and different types of TL calculations. Collectively these changes brought BELLHOP3D to a level of maturity approaching the original 2D BELLHOP. The Office of Naval Research has provided important additional support for studying out-of-plane effects and for further testing and development. Laurel Henderson implemented Ding Lee's FOR3D code at HLS and ran all the test cases that used FOR3D. Xavier Zabal provided the analytic solutions for the perfect wedge, the truncated wedge, and the seamount.```

